GLACIAL CYCLES: ROLE OF ECCENTRICITY IN ICE LINE MOVEMENT MATH CLIMATE SEMINAR 03/08/2016

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MILANKOVITCH CYCLES

Over the past one million years, glacial/interglacial cycles have occurred with periodicity of about 100,000 years.

Variations in the Earth's orbital parameters (obliquity, eccentricity, and precession) pace the glacial cycles.

COUPLED TEMPERATURE-ICE LINE MODEL

Time-dependent Energy Balance Model by Budyko:

$$R \frac{\partial T}{\partial t}(y,t) = Qs(y)(1 - \alpha(y)) - (A + BT) - C(T - \overline{T}).$$
 (1)
 $Qs(y)(1 - \alpha(y)) =$ absorbed term
 $(A + BT) =$ emitted term.
 $C(T - \overline{T}) =$ transport term.

The albedo function is defined as:

$$\alpha_{\eta}(y) = \begin{cases} \alpha_1, & \text{if } y < \eta \\ \alpha_2, & \text{if } y > \eta, \end{cases}$$

where $\alpha_1 < \alpha_2$.

COUPLED TEMPERATURE-ICE LINE MODEL

Couple Budyko's equation with Widiasih's ODE in [4] for the evolution of the ice line, η :

$$\frac{d\eta}{dt} = \rho(T(\eta, t) - T_c). \quad (2)$$

 T_c is a critical temperature above which ice melts and below which ice forms.

QUADRATIC APPROXIMATION

The above infinite dimensional system ((1) and (2)) is approximated by the system of ODEs as done by McGehee and Widiaish in [1]:

$$\begin{cases} \dot{w} = -\tau \left(w - F(\eta) \right) \\ \dot{\eta} = \rho \left(w - G(\eta) \right) \end{cases}$$
(3)

Where w, is a translate of the global average temperature. $F(\eta)$ (cubic polynomial) and $G(\eta)$ (quadratic polynomial) are given below:

$$F(\eta) = \frac{1}{B} \left(Q(1 - \alpha_0) - A + CL(\alpha_2 - \alpha_1) \left(\eta - \frac{1}{2} + s_2 P_2(\eta) \right) \right),$$

$$G(\eta) = -Ls_2(1 - \alpha_0)p_2(\eta) + T_c.$$

QUADRATIC APPROXIMATION

In [5], McGehee and Lehman proved that: $Q = Q(e) = \frac{Q_0}{\sqrt{1-e^2}},$ and in [2], McGehee and Widiasih showed that: $s_2(\beta) = \frac{5}{16}(-2+3\sin^2\beta).$

From [1]:

$$L = \frac{Q}{B + C},$$

$$P_{2}(\eta) = \frac{1}{2}(\eta^{3} - \eta),$$

$$p_{2}(\eta) = \frac{1}{2}(3\eta^{2} - 1).$$

QUADRATIC APPROXIMATION

Parameter	Value	Units
Q ₀	343	$W m^{-2}$
A	202	$W m^{-2}$
В	1.9	$W m^{-2} K^{-1}$
С	3.04	$W m^{-2} K^{-1}$
$lpha_1$	0.32	dimensionless
α_2	0.62	dimensionless
$\alpha_0 = \frac{\alpha_1 + \alpha_2}{2}$	0.47	dimensionless
T_{c}	-5.5 and -10	°C

EQUILIBRIUM SOLUTIONS

$$\begin{cases} \dot{w} = -\tau \big(w - F(\eta) \big) \\ \dot{\eta} = \rho \big(w - G(\eta) \big) \end{cases}$$

In order to find equilibrium solutions of the system above, we set the derivatives to 0, and we get $w = F(\eta) = G(\eta).$

Now, we may solve for η in the equation $F(\eta) - G(\eta) = 0.$

Three roots are found, and one is discarded since it doesn't belong to the range [0,1].

MILANKOVITCH FORCING AND ICE LINES At $T = -10^{\circ}$ C



MILANKOVITCH FORCING AND ICE LINES At $T = -5.5^{\circ}$ C



POWER SPECTRUM $T_c = -10^{\circ}$ C (SINK - SMALLER ICE CAP)



POWER SPECTRUM $T_c = -5.5^{\circ}C$ (SINK - LARGER ICE CAP)



REVISIT MODEL

The infinite dimensional system ((1) and (2)) is approximated by the system of ODEs as done by McGehee and Widiaish in [1]:

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ADDITION OF SNOW LINE



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 $B = \{(w, \eta, \xi) : w \in \mathbb{R}, \eta \in [0, 1], \xi \in [0, 1]\},\$ $b_0 < b < b_1$ = ablation rates, a =accumulation rate. When $b(\eta - \xi) - a(1 - \eta) < 0$, set $T_c = -5.5$ °C and $\begin{cases} \dot{w} = -\tau(w - F(\eta)) \\ \dot{\eta} = \rho(w - G_{-}(\eta)) \\ \dot{\xi} = \epsilon (b_0(\eta - \xi) - a(1 - \eta)). \end{cases}$ (4)When $b(\eta - \xi) - a(1 - \eta) > 0$, set $T_c = -10$ °C and $\begin{cases} \dot{w} = -\tau(w - F(\eta)) \\ \dot{\eta} = \rho(w - G_+(\eta)) \\ \dot{\xi} = \epsilon (b_1(\eta - \xi) - a(1 - \eta)). \end{cases}$ (5)

ADDITION OF SNOW LINE

We thus arrive at a 3-dimensional system having a plane of discontinuity [3]:

$$\Sigma = \{ (w, \eta, \xi) : b(\eta - \xi) - a(1 - \eta) = 0 \} \\= \{ (w, \eta, \xi) : \xi = \left(1 + \frac{a}{b} \right) \eta - \frac{a}{b} \}.$$

Constants	Value
а	1.05
b_0	1.5
b	1.75
b_1	5



RECALL

When
$$b(\eta - \xi) - a(1 - \eta) < 0$$
, set $T_c = -5.5$ °C and

$$\begin{cases}
\dot{w} = -\tau(w - F(\eta)) \\
\dot{\eta} = \rho(w - G_-(\eta)) \\
\dot{\xi} = \epsilon (b_0(\eta - \xi) - a(1 - \eta)).
\end{cases}$$



RECALL

When
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\end{cases}$$





RECALL

$$B = \{(w, \eta, \xi) : w \in \mathbb{R}, \eta \in [0,1], \xi \in [0,1]\}, \\ b_0 < b < b_1 = \text{ablation rates,} \\ a = \text{accumulation rate.} \\ \text{When } b(\eta - \xi) - a(1 - \eta) < 0, \text{ set } T_c = -5.5 \text{ °C and} \\ \begin{cases} \dot{w} = -\tau(w - F(\eta)) \\ \dot{\eta} = \rho(w - G_-(\eta)) \\ \dot{\xi} = \epsilon (b_0(\eta - \xi) - a(1 - \eta)). \end{cases}$$
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DISCUSSION AND FURTHER WORK

Use a more complex model backed by intensive numerical simulations to verify the role eccentricity plays.

Why is η_{saddle} in the $T = -5.5^{\circ}\text{C}$ case perfectly in-sync with the obliquity variation whereas η_{sink} is forced by eccentricity?

Long shot: Relate to the Mid-Pliestocene Transition – shift from 41 kYr (obliquity forced) cycles to 100 kYr cycles

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